

Strongly-coupled Anisotropic Gauge Theories And Holography

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Based on works with: U. Gursoy(Utrecht Univ.) and J. Pedraza(Univ. of Amsterdam); arxiv:1708.05691

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Outline

- 1 Introduction
- 2 The Theory
- 3 Phase Transitions
- 4 Conclusions

Short talk: Briefly we have

- Strongly coupled **anisotropic** theory.
- How the theory looks like and how to obtain it?
 - ✓ 4d $SU(N)$ gauge theory in the large N_c -limit.
 - ✓ Its dynamics are affected by a **scalar operator** $\mathcal{O} \sim \text{Tr} F^2$.
 - ✓ Anisotropy is introduced by **another operator** $\tilde{\mathcal{O}} \sim \theta(x_3) \text{Tr} F \wedge F$ with a space dependent coupling.
 - ✓ On the gravity dual side we have a "backreacting" scalar field depending on spatial directions, the **axion**; and a non-trivial **dilaton**.
- Eventually the gravity dual theory is an **Einstein-Axion-Dilaton theory** in 5 dimensions with a non-trivial potential.
 - ✓ Solutions are RG flows:
AdS in UV \Rightarrow **Anisotropic (Hyperscaling Lifshitz-like)** in IR.

Theory Evolution:

Non-Confining:

(Azeyanagi, Li, Takayanagi, 2009; Mateos, Trancanelli, 2011; Jain, Kundu, Sen, Sinha, Trivedi, 2015;...)

Confining:

(D.G., Gursoy, Pedraza, 2017)

Similar ideas in different context:

(Gaiotto, Witten 2008; Chu, Ho, 2006; Choi, Fernandez, Sugimoto 2017;...)

Instead of motivation:

- Such theories **violate** well known universal relations:
 - ✓ **Shear Viscosity over entropy density ratio** is parametrically **lower** than KSS bound: $1/4\pi$.
(Rebhan,Steineder 2011; Jain, Samanta, Trivedy 2015; D.G., Gursoy, Pedraza, 2017;...)
 - ✓ **Langevin Coefficient inequality for heavy quark motion** in the **anisotropic** plasma $\kappa_L > <\kappa_T$.
(Gursoy, Kiritsis, Mazzanti, Nitti 2010; D.G, Soltanpanahi, 2013a, 2013b)
 - ✓ ...

Especially interesting since several **strongly coupled systems** exhibit **anisotropies**.

The Anisotropic Theory

The generalized **Einstein-Axion-Dilaton action** with a **potential** for the dilaton and an **arbitrary coupling** between the axion and the dilaton:

$$S = \frac{1}{2\kappa^2} \int d^5x \sqrt{-g} \left[R - \frac{1}{2}(\partial\phi)^2 + V(\phi) - \frac{1}{2}Z(\phi)(\partial\chi)^2 \right].$$

The eoms read

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{1}{2}\partial_\mu\phi\partial_\nu\phi + \frac{1}{2}Z(\phi)\partial_\mu\chi\partial_\nu\chi - \frac{1}{4}g_{\mu\nu}(\partial\phi)^2 - \frac{1}{4}g_{\mu\nu}Z(\phi)(\partial\chi)^2 + \frac{1}{2}g_{\mu\nu}V(\phi),$$

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu\phi) = \frac{1}{2}\partial_\phi Z(\phi)(\partial\chi)^2 - V'(\phi),$$

$$\frac{1}{\sqrt{-g}}\partial_\mu(\sqrt{-g}g^{\mu\nu}\partial_\nu\chi) = 0.$$

Where the functions

$$V(\phi) = 12 \cosh(\sigma\phi) + \left(\frac{-m(\Delta)^2}{2} - 6\sigma^2 \right) \phi^2, \quad Z(\phi) = e^{2\gamma\phi},$$

((Gubser, Nellore), Pufu, Rocha 2008a,b)

Note: For $\sigma = 0, \gamma = 1, m(\Delta) = 0$ the action and the solution of eoms, are of IIB supergravity.

(Mateos, Trancanelli, 2011)

A Solution: Schematically

- The **background solution**

$$ds^2 = \frac{e^{-\frac{1}{2}\phi(u)}}{u^2} \left(-\mathcal{FB} dt^2 + dx_1^2 + dx_2^2 + \mathcal{H} dx_3^2 + \frac{du^2}{\mathcal{F}} \right),$$

$$\chi = \alpha x_3, \quad \phi = \phi(u),$$

- Note:**

The **linear axion** simplifies **tremendously** the system of equations!

- We **solve** the system of eoms to obtain the theory e.g. for $\Delta = 4$: Geometries that are **AdS in UV** flowing to **Hyperscaling Lifshitz-like violation geometries in IR** with arbitrary scalings:

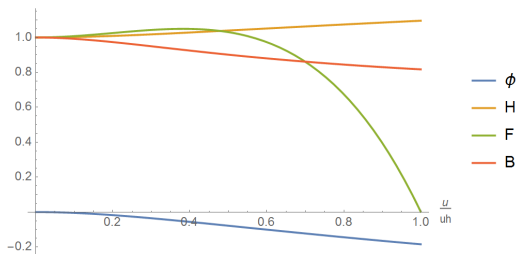
$$ds^2 = u^{-\frac{2\theta}{3}} \left(-u^{2z} (f(u) dt^2 + dx_{1,2}^2) + \tilde{\alpha} u^2 dx_3^2 + \frac{du^2}{f(u) u^2} \right),$$

- Note:** The values of (θ, z) depend on Axion-Dilaton coupling and Potential parameters (σ, γ) . In supergravity there is a **single** solution $(\theta = 0, z = 3/2)$. (*Azeyanagi, Li, Takayanagi, 2009*)

Solution : The Full Flow

- Fixing (γ, σ) and α and u_h we get the **metric flow** from boundary to horizon:

$$ds^2 = \frac{e^{-\frac{1}{2}\phi(u)}}{u^2} \left(-\mathcal{F}B dt^2 + dx_1^2 + dx_2^2 + \mathcal{H}dx_3^2 + \frac{du^2}{\mathcal{F}} \right),$$



- In sufficiently **high temperatures**, $T \gg \alpha$ the Einstein equations can be solved analytically; e.g. for $\gamma = 1, \sigma = 0$:

$$\mathcal{F}(u) = 1 - \frac{u^4}{u_h^4} + \alpha^2 \frac{1}{24u_h^2} \left[8u^2(u_h^2 - u^2) - 10u^4 \log 2 + (3u_h^4 + 7u^4) \log \left(1 + \frac{u^2}{u_h^2} \right) \right]$$

$$\mathcal{B}(u) = 1 - \alpha^2 \frac{u_h^2}{24} \left[\frac{10u^2}{u_h^2 + u^2} + \log \left(1 + \frac{u^2}{u_h^2} \right) \right], \quad \mathcal{H}(u) = \left(1 + \frac{u^2}{u_h^2} \right)^{\frac{\alpha^2 u_h^2}{4}}.$$

We have obtained the theories, are all of them **physical**
and **stable**?

Local Thermodynamical Stability in CE and the request for physical theories (NEC) constrain the parameters:

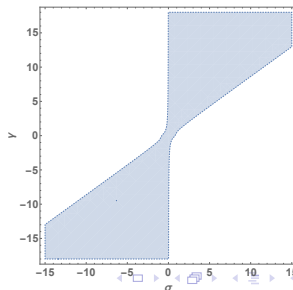
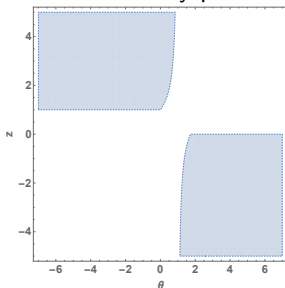
Three conditions that constrain (z, θ) and as a result (γ, σ) .

$$(z - 1)(1 - \theta + 3z) \geq 0, \quad z = \frac{2 + 4\gamma^2 - 3\sigma^2}{2\gamma(2\gamma - 3\sigma)},$$

$$\theta^2 - 3 + 3z(1 - \theta) \geq 0, \quad \theta = \frac{3\sigma}{2\gamma},$$

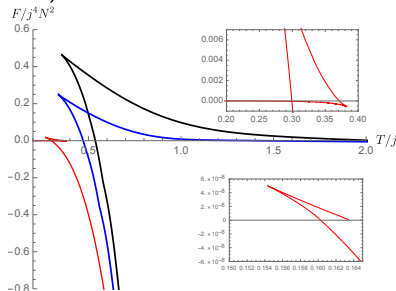
$$1 - \theta + 2z \geq 0.$$

Translated to the following diagrams where the blue region is the acceptable for the theory parameters.



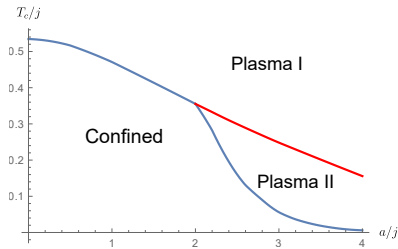
Confinement/Deconfinement Phase transitions

- The free energy of the theories vs the temperature T for different anisotropy ($\alpha/j=0,1,3$):



- Horizontal Axis: **Confining Phase.**
- Upper Branch: **Black hole A:Deconfining Plasma Phase.**
- Lower Branch: **Black hole B:Deconfining Plasma Phase.**
- $\alpha/j \simeq 2$: A critical value above which a richer structure in the phase diagram exist.

- The **Critical Temperature** of the theories vs the **anisotropy** gives:



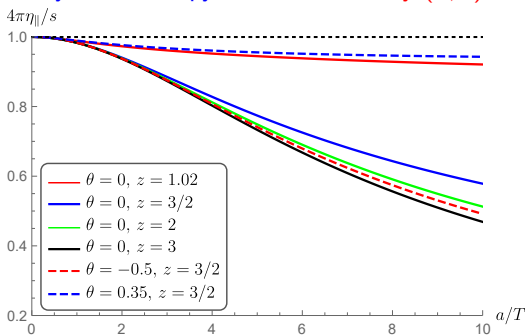
- The T_c is **reduced** in presence of anisotropies of the theory.

The Proposal

- The $T_c(\alpha)$ decrease with α , resembling the phenomenon of **inverse magnetic catalysis** where the **confinement-deconfinement** temperature decreases with the magnetic field B (where an anisotropy is introduced as in our plasma).
- **No charged fermionic degrees** of freedom in our case; our plasma is neutral.
- Our findings suggest that the **anisotropy by itself** could instead be the cause of lower T_c in presence of **anisotropies**.

η/s for our theory: Dependence on the Temperature.

The shear viscosity over entropy ratio for arbitrary (z, θ) .



- The ratio depends on the temperature at $\alpha/T \gg 1$ as

$$\frac{\eta_{||}}{s} \sim \left(\frac{T}{\tilde{\alpha}|1+3z-\theta|} \right)^{2-\frac{2}{z}}.$$

- The range of the temperature power is $[0, \infty)$.

Conclusions

- ✓ We have presented **Confining Anisotropic** theories and studied the **Confinement/Deconfinement** phase transitions.
- ✓ The theories are **physical and stable** for a wide range of **parameters** of the theory.
- ✓ The **Confinement/Deconfinement** phase transitions occur at **lower** critical **Temperature** as the anisotropy is **increased**!
- ✓ We suggest that the **anisotropy by itself** could instead be a cause of **inverse magnetic catalysis**.
- ✓ The **shear viscosity over entropy density** ratio, takes values parametrically **lower** than $1/4\pi$, and depends on the **Temperature** as $T^{2-2/z}$.
- ✓ The **diffusion (butterfly velocity) of chaos** occurs **faster** than isotropic systems.
- More studies needed on such theories...

Thank you!

Null Energy Condition

- The averaged radial acceleration between two null geodesics is

$$A_r = -4\pi T_{\mu\nu} N^\mu N^\nu ,$$

if it is negative the null geodesics observe a **non-repulsive gravity** on nearby particles along them.

- This imposes the **Null Energy Condition**

$$T_{\mu\nu} N^\mu N^\nu \geq 0 , \quad N^\mu N_\mu = 0 ,$$

leading to the following constraints:

- For the **Lifshitz-like** space $z \geq 1$.
- For the **Hyperscaling violation anisotropic metric** in 3+1-dim spacetime and anisotropic in 1-dim reads

$$(z-1)(1-\theta+3z) \geq 0 ,$$
$$\theta^2 - 3 + 3z(1-\theta) \geq 0 .$$

Additional conditions from **thermodynamics**?

Local Thermodynamic Stability

- The **necessary and sufficient conditions** for **local thermodynamical stability** in the **canonical ensemble** are

$$c_\alpha = T \left(\frac{\partial S}{\partial T} \right)_\alpha \geq 0, \quad \Phi' = \left(\frac{\partial \Phi}{\partial \alpha} \right)_T \geq 0$$

c_α is the **specific heat**: increase of the temperature leads to increase of the entropy.

Φ' is **derivative of the potential**: the system is stable under infinitesimal charge fluctuations.

- In the **GCE** these conditions should be equivalent of having **no positive eigenvalues** of the **Hessian matrix** of the entropy with respect to the thermodynamic variables. *(Gubser, Mitra 2001)*
- In the IR the **positivity** of **the specific heat** imposes

$$c_\alpha = 1 - \theta + 2z \geq 0$$